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**1 A breaking internal wave in the surface ocean**  
**2 boundary layer**

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3 **Abstract.** High-temporal resolution measurements in the Labrador Sea  
4 surface layer are presented using an upwardly-profiling autonomous microstruc-  
5 ture instrument, which captures an internal wave in the act of breaking at  
6 the base of the surface mixed layer, driving turbulence levels two to three  
7 orders of magnitude above the background. While lower-frequency (near-inertial)  
8 internal waves are known to be important sources of turbulence, we report  
9 here a higher frequency internal wave breaking near the ocean surface. Due  
10 to observational limitations, the exact nature of the instability cannot be con-  
11 clusively identified, but the interaction of wave-induced velocity with unre-  
12 solved background shear appears to be the most likely candidate. These ob-  
13 servations add a new process to the list of those currently being considered  
14 as potentially important for near-surface mixing. The geographical distribu-  
15 tion and global significance of such features is unknown, and underscores the  
16 need for more extensive small-scale, rapid observations of the ocean surface  
17 layer.

## 1. Introduction

18 The ocean surface mixed layer is a critical interface for air-sea interaction. The physical  
19 processes controlling turbulent mixing within the ocean surface mixed layer impact the  
20 air-sea exchange of momentum, heat, and carbon dioxide, and also set the properties  
21 of ventilated water masses that eventually subduct into the ocean interior. For these  
22 reasons, global climate models are known to be sensitive to parameterizations of near-  
23 surface turbulent mixing [*Fox-Kemper et al.*, 2011; *Belcher et al.*, 2012], as this modulates  
24 the mixed layer depth [*Stevens et al.*, 2011].

25 Higher-frequency internal waves in the upper ocean have been shown to be a source  
26 of turbulence, mixing, and vertical heat fluxes through the pycnocline [*Barton et al.*,  
27 2001; *Dewey et al.*, 1999]. Away from topography, such waves can be generated in a  
28 variety of ways, most of which originate with input of wind energy into the mixed layer.  
29 Inertial motions can excite a wide spectrum of internal waves, from low mode near-inertial  
30 waves that can propagate far from their generation site to higher-frequency waves which  
31 dissipate locally [*Bell*, 1978; *Simmons and Alford*, 2012]. Perturbations on the base of  
32 mixed layer from Langmuir cells [*Polton et al.*, 2008; *Belcher et al.*, 2012], nighttime  
33 convection [*Wijesekera and Dillon*, 1991; *Ansong and Sutherland*, 2010], shear instabilities  
34 [*Lien et al.*, 2002; *Dohan and Davis*, 2011], and turbulent patches [*Dohan and Sutherland*,  
35 2003] can all generate internal waves with frequencies near the buoyancy frequency  $N$ .  
36 These waves can transfer energy from the mixed layer and through the transition layer  
37 —the highly stratified region at the base of the surface mixed layer— to the pycnocline,  
38 where they remain trapped because they cannot propagate through the weakly stratified

39 region below the pycnocline [*Johnston and Rudnick, 2009*]. These waves can then break  
40 and generate turbulence through self-induced shear or steepening through interactions  
41 with lower frequency waves or background shear [*Thorpe, 2005*]. The relative importance  
42 of these various processes is not yet known, with observational challenges being recognized  
43 as a limiting factor [*Belcher et al., 2012; Sutherland et al., 2014*].

44 Near-surface mixing processes take on a particular significance in the Labrador Sea, a  
45 region of the North Atlantic’s subpolar gyre that is one of the key locations for the Atlantic  
46 Meridional Overturning Circulation (AMOC). There, intense wintertime cooling leads to  
47 deep convection to depths of 500–2000 m, forming a mode water known as Labrador  
48 Sea Water [e.g. *Lilly et al., 1999*]. Each spring, a rapid surface freshening is observed  
49 across the Labrador Sea basin [*Lilly et al., 1999; Straneo, 2006; Schmidt and Send, 2007*],  
50 resulting in a 200–300 m thick highly buoyant layer that serves as a barrier to the next  
51 winter’s deep convection. The predictability of deep water ventilation in the Labrador  
52 Sea requires understanding not only of the variability in the wintertime surface heat loss,  
53 but also the near-surface processes that contribute to rebuilding a stable stratification  
54 [*Våge et al., 2009*]. While this rapid surface capping appears to be a generic feature of  
55 convective regions [*Marshall and Schott, 1999*], little is known about its dynamics. With  
56 freshwater discharge from the Arctic now dramatically increasing [*Peterson et al., 2006*], it  
57 is important to understand Labrador Sea surface processes at work during restratification.

58 Motivated by the importance of near-surface mixing in general, and in deep convection  
59 regions in particular, the first direct microstructure turbulence measurements from the  
60 Labrador Sea were collected during May 2010. These were obtained using the Air-Sea  
61 Interaction Profiler (ASIP), an autonomous, upwardly-profiling instrument [*Ward et al.,*

2014; *Sutherland et al.*, 2014; *Callaghan et al.*, 2014; *Sutherland et al.*, 2013]. This instrument is specifically designed to make high-resolution microstructure profiles of the ocean surface layer, a climatically important region of the ocean for which very few direct microstructure measurements are available. Other microstructure platforms typically profile downward starting from a depth of tens of meters, and focus on interior turbulence with temporal resolutions of 30 minutes or more. This dataset therefore provides a rare, high-resolution view into rapidly evolving near-surface dynamics.

The main result is the presentation of a breaking internal wave at the base of the Labrador Sea mixed layer. This adds a new candidate to the list of potential sources of upper-ocean turbulence, and offers a counterexample to other works [see e.g. *Johnston and Rudnick*, 2009] that identify near-inertial shear as the primary driver.

## 2. Observations

During May 2010, microstructure measurements of the Labrador Sea surface layer were collected as a part of a hydrographic cruise carried out annually by the Bedford Institute of Oceanography (BIO). Since 1990, BIO has repeatedly occupied the AR7W line of stations that stretches from the coast of Labrador to the coast of Greenland [e.g., *Lazier et al.*, 2002; *Yashayaev*, 2007; *Yashayaev and Loder*, 2009], as shown in Figure 1. Conductivity-temperature-depth (CTD) profiles of the upper water column along this section, see Figure 2, reveal a thin, 30–60 m buoyant layer over a largely unstratified interior. This structure is typical of the springtime Labrador Sea, a consequence of rapid freshwater restratification following deep convection [*Lilly et al.*, 1999; *Straneo*, 2006; *Schmidt and Send*, 2007].

83 A full-depth ( $\sim 3500$  m) CTD profile taken contemporaneously with the microstruc-  
84 ture measurements describes the hydrographic background, see Figure 3. The buoyancy  
85 frequency  $N$  was computed by first completing a piecewise fit to the observed potential  
86 density  $\sigma_0$  profile, consisting of 14 linear segments above an exponential decay from 230 m  
87 to the bottom. This approach seemed to offer the best balance between preserving sharp  
88 gradients, while still averaging over small-scale noise. An accurate estimation of  $N$  allows  
89 us to determine the depth range over which internal waves will be trapped.

90 This CTD profile reveals a relatively warm, fresh surface layer about 60 m thick, above  
91 a weakly stratified interior with a buoyancy frequency of about  $N = 0.66$  cycles per hour  
92 (cph), or one cycle per 90 minutes (Figure 3). The transition layer—the highly stratified  
93 region between the well-mixed surface layer and the deep stratification—was roughly 10 m  
94 thick, with a peak buoyancy frequency of  $N = 6.6$  cph, an order of magnitude higher than  
95 weak interior value. For comparison, the inertial frequency at this latitude is 0.07 cph, or  
96 one cycle per 14.4 hours.

97 In addition, a shipboard Acoustic Doppler Current Profiler (ADCP), a Teledyne RDI  
98 Ocean Surveyor II 75kHz, recorded five minute averages of horizontal velocities in 8 m  
99 bins between 20 m and 650 m depth, see Figure 3b. Even taking the mean across the four  
100 hour time series, we found that 40 m averaging in the vertical was necessary to reduce  
101 noise. As a result of this temporal and vertical averaging, the resolvable shear is quite  
102 weak, around  $0.02 \text{ m s}^{-1}$  per 100 m over the depth interval containing the transition layer.

103 The microstructure data was acquired with ASIP, a 2.5 m long upward-profiling instru-  
104 ment that is outfitted with two shear probes, fast response conductivity and temperature  
105 (C/T) microstructure sensors, and slower more accurate C/T sensors. ASIP is unteth-

106 ered and profiles autonomously from as deep as 100 m to the sea surface. The instrument  
107 submerges itself to a preset depth using thrusters, and subsequently rises under its own  
108 buoyancy until it penetrates the surface. The rapid profiling capability of ASIP allows for  
109 processes with short time scales to be observed [see also *Sutherland et al.*, 2013].

110 On May 22, 2010, in the vicinity of the hydrographic profile shown in Figure 3, a 4-hour  
111 microstructure dataset was acquired using ASIP. During the deployment, ASIP drifted at  
112 about  $0.15 \text{ m s}^{-1}$  to the southeast, see Figure 4. The dataset consisted of 30 profiles of  
113 high-resolution temperature, conductivity, and velocity shear from 100 m to the surface,  
114 with approximately eight minutes between each profile, as presented in Figure 5.

115 The temperature and conductivity data from ASIP were averaged in half-meter verti-  
116 cal bins, after which salinity, density, and buoyancy frequency were computed following  
117 *McDougall and Barker* [2011]. Profiles were smoothed with 10 m running means before  
118 gradients were computed, for example, in the calculation of the buoyancy frequency pre-  
119 sented in the upper panel of Figure 5. The mean density profile from ASIP is shown in  
120 Figure 3b for comparison with the CTD cast. The pycnocline is sharper in the CTD cast,  
121 as the 10-m amplitude oscillations in the pycnocline (discussed in detail below) broaden  
122 the density gradient in the ASIP profiles upon averaging.

123 The turbulent dissipation  $\epsilon$  was estimated from the shear power spectra using standard  
124 methods [*Yamazaki and Osborn*, 1990; *Oakey*, 1982]. The spectra were calculated over  
125 windows of 1000 points in length, corresponding to 0.5 s of sampling. The windows  
126 were shifted in 500 point intervals ensuring 50% overlap between adjacent  $\epsilon$  estimates,  
127 resulting in 0.25 m vertical bins for dissipation rate. Following *Osborn* [1980], the vertical  
128 eddy diffusivity  $K_\rho$  was then estimated from turbulent dissipation by assuming a balance

129 between mechanical production, buoyancy flux, and turbulent dissipation in the turbulent  
 130 kinetic energy equation. This leads to a vertical eddy diffusivity given by  $K_\rho = \Gamma\epsilon/N^2$ ,  
 131 where the coefficient  $\Gamma$  is a mixing efficiency that is conventionally taken to be 0.2.

### 3. Results and Discussion

132 The displacement of the pycnocline seen in Figure 5a clearly reveals the presence of  
 133 an internal wave, with a frequency of about 1.2 cph, or a 50 minute period, and large-  
 134 amplitude isopycnal excursions of about 10 m in the vertical. Associated with these  
 135 displacements is a patch of elevated turbulent dissipation at the base of the mixed layer  
 136 between 45 and 55 m depth, indicated by the boxes in both panels of Figure 5, within  
 137 which the dissipation rate  $\epsilon$  reached values as high as  $2 \times 10^{-7} \text{ m}^2 \text{ s}^{-3}$ . The deep patch of  
 138 elevated dissipation is separated from the surface dissipation by a  $\sim 20$  m thick layer of very  
 139 weak dissipation, with  $\epsilon$  values as low as  $\epsilon \sim 2 \times 10^{-10} \text{ m}^2 \text{ s}^{-3}$ . Therefore it appears clear  
 140 that this patch is associated with the internal wave itself, rather than with turbulence-  
 141 generating mechanisms at the surface. These enhanced dissipation rates approach the  
 142 values associated with near-surface mixing, with peak values of  $\epsilon \sim 1 \times 10^{-6} \text{ m}^2 \text{ s}^{-3}$ ,  
 143 which are driven by surface heat losses and extend to roughly 30 m depth.

144 Until 19:44 UTC (Figure 5), the depth-averaged dissipation rate near the base of the  
 145 mixed layer (between depths of 45 and 55 m) was low, averaging  $7 \times 10^{-10} \text{ m}^2 \text{ s}^{-3}$  over  
 146 the first 10 profiles, which corresponds to a vertical turbulent eddy diffusivity of  $6 \times$   
 147  $10^{-6} \text{ m}^2 \text{ s}^{-1}$ . Between 19:44 and 21:02, during the patch of elevated turbulence, the  
 148 depth-averaged dissipation rate exceeded  $2 \times 10^{-8} \text{ m}^2 \text{ s}^{-3}$ , with an average over these 10  
 149 profiles of  $1 \times 10^{-8} \text{ m}^2 \text{ s}^{-3}$ . The eddy diffusivity computed from the averaged dissipation  
 150 profile during this time was elevated to  $1 \times 10^{-4} \text{ m}^2 \text{ s}^{-1}$ , with a maximum value of

151  $6 \times 10^{-4} \text{ m}^2 \text{ s}^{-1}$ , two orders of magnitude above the background values. After 21:02, the  
152 mean depth-averaged dissipation rate of the last 10 profiles dropped to  $3 \times 10^{-9} \text{ m}^2 \text{ s}^{-3}$ ,  
153 which was still elevated above the background values, and which appeared at that time  
154 to detach vertically from the pycnocline base. At this same time, the turbulence from  
155 the surface begins penetrating to the depth of the pycnocline, perhaps due to convective  
156 overturns which lead to the observed statically unstable density field.

157 Wave speed and wavelength were estimated in two different ways, using different as-  
158 sumptions, both giving similar results. The first approach attempts to match the observed  
159 vertical distribution of isopycnal displacements with a suitable vertical mode. Vertical  
160 structure functions (which are independent of frequency) for the vertical modes were  
161 computed from the full-depth density profile using standard methods [e.g. *Gill*, 1982],  
162 and the upper 600 m of these structure functions are shown in Figure 3d. The shallowest  
163 extremum in vertical displacement amplitude for each mode was at 1892 m, 823 m, 295 m,  
164 104 m, and 67 m, for modes 1–5, respectively. The internal wave captured by ASIP has  
165 a peak vertical displacement amplitude at 70 m, which is best matched by the extremum  
166 in mode five. The wave speed for this mode is  $0.28 \text{ m s}^{-1}$ ; for the observed frequency  
167 of wave, this corresponds to a horizontal wavenumber  $k = \omega/c = 0.0075 \text{ rad m}^{-1}$  or a  
168 horizontal wavelength of 840 m.

169 Alternatively, assuming that the wavelength is long compared to the depth of the surface  
170 layer, we can idealise the observed internal wave as an interfacial wave between a shallow  
171 surface layer and an infinitely deep lower layer. The long wave approximation can then  
172 be used to estimate the wave speed [see *Kundu and Cohen*, 2002] as  $c = \sqrt{g'H}$ , where  
173  $g'$  is the reduced gravity  $g' = (\Delta\rho/\rho_o)g$ ,  $\Delta\rho$  is the density change over the interface,

174  $\rho_o$  is a reference density of  $1027 \text{ kg m}^{-3}$ , and  $H$  is the thickness of the surface layer,  
175 which here was chosen to be 60 m based on the average mixed layer depth from the  
176 ASIP measurements. In this calculation,  $\Delta\rho$  was computed as the difference between  
177  $\sigma_0$  above and below the pycnocline. From the measurements,  $g'$  was found to be about  
178  $0.002 \text{ m s}^{-2}$ , again yielding a phase velocity of  $0.3 \text{ m s}^{-1}$ ; thus both approaches lead to  
179 the same approximate estimate of the wave speed and wavelength of the waves observed  
180 here.

181 Due to the Labrador Sea's weak interior stratification, a wave of this frequency will be  
182 strongly trapped in the upper water column. The white lines in Figure 2 mark a buoyancy  
183 frequency of  $N = 1.2 \text{ cph}$ , indicating that the wave cannot exist below about 200 m at the  
184 ASIP measurement location. To further examine this trapping, the results of a ray tracing  
185 calculation are also shown in Figure 2. This ray tracing considered only  $N$  and not velocity  
186 shear, a standard approach which should be a good approximation in the Labrador Sea  
187 interior, away from strong currents. The weak stratification in the Labrador Sea basin  
188 after convection acts as a waveguide, confining waves at this frequency to the uppermost  
189 100–300 m over this 300 km section. Thus the observed waves must have been generated  
190 by surface forcing of some type, as opposed to a deep source such as tidal interaction  
191 with bathymetry. Because of the strong waveguide, the horizontal location of the wave's  
192 origin cannot be identified, and in particular, a nonlocal origin is a distinct possibility.  
193 But the observed turbulent event here implies that horizontal propagation will be limited  
194 as energy is drained from the internal wave during breaking.

195 Based on the estimates above, we can attempt to use the available data to shed light  
196 on possible breaking mechanisms. A number of process could conceivably lead to the

197 breaking of an internal wave. An internal wave itself can generate convective instabilities  
 198 or shear instabilities that lead to breaking. For a monochromatic wave in a fluid with  
 199 an interface of finite thickness, *Fringer and Street* [2003] numerically showed how the  
 200 required steepness for breaking depends upon the non-dimensional interface thickness  $\delta k$ ,  
 201 where  $\delta$  is a measure of the interface thickness defined in *Fringer and Street* [2003] and  
 202  $k$  is the horizontal wavenumber. The amplitude  $a$  of the waves here is approximately 10  
 203 m and  $\delta$  is 35 m. For the wave parameters estimated above, the  $ak$  is 0.075 and  $\delta k$  is  
 204 0.26. *Fringer and Street* [2003] found convective overturns for  $\delta k$  greater than 2.33 and  
 205  $ak$  greater than 1, thus such instabilities are not likely to play a role here.

206 Below  $\delta k = 2.33$ , shear instabilities were the observed breaking mechanism; the bottom  
 207 end of the parameter range in *Fringer and Street* [2003] was  $\delta k = 0.31$ . The labora-  
 208 tory experiments of *Troy and Koseff* [2005] investigate breaking interfacial waves at non-  
 209 dimensional thicknesses smaller than this, which are typical of longer waves, as observed  
 210 here. They found that breaking occurred at a critical wave steepness  $ak_c = \sqrt{(2\delta k)}$ , where  
 211 here  $ak_c$  is 0.72, an order of magnitude larger than the measured  $ak$ . Thus the waves are  
 212 not inherently steep enough to break without the presence of shear from other sources  
 213 [*Thorpe*, 1978].

214 Interaction with background shear could lead to wave breaking. Between 44 and 55 m,  
 215 where the breaking occurs,  $N = 0.0042$  rad s<sup>-1</sup>; the change in velocity over this depth  
 216 necessary to reduce the Richardson number below 1 is 0.046 m s<sup>-1</sup>. To first order, the  
 217 wave-induced velocity  $u_w$  can be approximated as  $a\omega$  [*Thorpe*, 2005]. The amplitude here  
 218 is approximately 10 m, leading to an estimate of  $u_w = 0.021$  m s<sup>-1</sup>. If the shear used  
 219 in computing the Richardson number is the sum of the background shear and the wave

220 induced shear ( $du/dz \sim \Delta U/\delta + a\omega/\delta$ ), this wave-induced velocity enhanced by modest  
221 background shear would be sufficient to induce breaking; the velocity difference needed  
222 here is about  $0.025 \text{ m s}^{-1}$ .

223 Shipboard ADCP data is available during the time of the ASIP deployment. Thirty  
224 minute average ADCP velocity profiles are shown in Figure 6 without any vertical smooth-  
225 ing. The velocities in the surface mixed layer ( $0.15\text{--}0.20 \text{ m s}^{-1}$ ) are about an order of  
226 magnitude larger than the required velocity difference. The magnitude of variance is high,  
227 and is of the same order as the velocity difference needed to induce turbulence. However,  
228 after carrying out the processing that is necessary to reduce noise, with a five minute  
229 temporal average and 40 m smoothing, we find far lower values of the shear. There is,  
230 however, an observed acceleration of background environment by  $0.05\text{--}0.1 \text{ m/s}$  at the time  
231 of the elevated turbulence, suggesting that an unresolved change in the shear at this time  
232 is not implausible.

233 Thus the physical mechanism responsible for the elevated turbulence levels cannot be  
234 conclusively identified using the available data. Based on theoretical and laboratory based  
235 parameterisations, instabilities are not predicted, but the isopycnals appear to be vertical  
236 at the crest of the wave (Figure 5b). A reasonable explanation is that the background  
237 shear is under-represented by the shipboard ADCP measurements, as these data are  
238 more limited in their resolution than are ASIP's high-resolution measurements of the  
239 wave properties. This blurring of vertical gradients due to the averaging employed could  
240 have also obscured the required shear for generating Kelvin-Helmholtz instabilities (or  
241 potentially Holmboe instabilities if the maximum shear and density gradient are not co-  
242 located, e.g. *Smyth and Winters* [2003]). This difficulty is not unique to our dataset;

243 other authors have faced similar difficulties in identifying the source of shear instabilities  
244 using ADCP velocity data [Moum *et al.*, 2003].

245 Because of the issues surrounding estimating small scale velocity shear with shipboard  
246 ADCPs, we used the isopycnal displacements as measured by ASIP to estimate the fine-  
247 scale velocity structure following Moum *et al.* [2003]. In this approach, the velocity profile  
248 is inferred from the density profile by assuming that in the wave's frame of reference that  
249 the streamlines are parallel to the isopycnals and thus the velocity along each isopycnal  
250 can be estimated. See Moum *et al.* [2003] for further details. The isopycnal separation  
251 ( $\Delta\rho = 0.04 \text{ kg m}^{-3}$  in the current dataset) was chosen to remove fluctuations due to  
252 turbulence so that the remaining density profile is composed of the background profile  
253 and the wave. While coarsely resolved, the background velocity profile was taken to be  
254 the average profile from the ADCP and was relatively constant within the upper 100 m  
255 ( $0.18 - 0.20 \text{ m s}^{-1}$ ). The background density profile was taken as the mean density profile  
256 from the deployment. This method is valid until the turbulent event, after which it cannot  
257 be determined if the straining of isopycnals is from the wave or from mixing.

258 Using the highlighted isopycnals in Figure 5, the fine scale velocity structure was inferred  
259 for the first ten profiles of the deployment. The velocity profile from Profile 10 (19:43) is  
260 shown in Figure 7, where the isopycnals are compressed relative to the background profile,  
261 generating a shear layer in the the pycnocline. Note that the irregular spacing between  
262 data points is due to the use of isopycnal coordinates for this calculation. The sharp drop  
263 at the base of the mixed layer does appear to be a real feature, and is comparable to those  
264 observed in Moum *et al.* [2003] Figure 18.

265 From the velocities estimated here along the highlighted isopycnals, the Richardson  
266 number between the 2nd and 3rd isopycnals (where the turbulence is initiated) drops  
267 below 0.1 in the trough of the wave at Profile 8 (19:29) where the first observations of  
268 elevated turbulence are seen (Figure 5). The Richardson number then drops below 0.01  
269 at Profile 10. This satisfies the more stringent Richardson number criterion of *Barad*  
270 *and Fringer* [2010], although with the existing dataset we are unable to determine if the  
271 length of time that the fluid is subjected to this low Richardson number is consistent  
272 with the growth rate of the instabilities. This has been found to play an important role  
273 in the breaking of interfacial waves in the laboratory [*Troy and Koseff*, 2005; *Barad and*  
274 *Fringer*, 2010]. While there are many assumptions in the approach taken here, the results  
275 indicate that finescale velocity from the wave passing through the ambient velocity field  
276 might be sufficient to initiate instability and thus turbulence. Once the turbulence is  
277 initiated, there is clear straining of the isopycnals after the crest of the wave, which has  
278 been associated with turbulent overturns in other environments [*Alford and Pinkel*, 2000]  
279 and thus might contribute to sustaining the turbulence.

280 While oscillations of the mixed layer base are commonly observed, well-resolved mea-  
281 surements of such higher-frequency internal waves are rather rare. One study that has  
282 previously reported on a similar wave, *Joyce and Stalcup* [1984], used acoustic obser-  
283 vations to identify a 10 m amplitude, 10 minute period oscillation of the stratification  
284 below the mixed layer; their Figure 5 appears strikingly similar to the upper panel of our  
285 Figure 5. The stratification in that study, within a warm core ring, was described as con-  
286 sisting of a buoyant surface layer above a homogeneous water mass, separated by a strong

287 thermocline at 45 m. Thus these observations occurred in a similar surface-intensified  
288 stratification to that of the post-convection Labrador Sea.

289 An outstanding mystery is the mechanism or mechanisms responsible for generating  
290 these waves. As described in the introduction, there are several possible mechanisms that  
291 could account for local generation of high-frequency waves, but it is not clear which if  
292 any of these is the most likely. Earlier in the cruise, the central Labrador Sea experienced  
293 several days of strong winds (Figure 8) and attendant long-period swell, an energetic  
294 environment that would seem readily capable of giving rise to internal waves. In a recent  
295 review paper, *Polzin and L'vov* [2011, see their Section 5.3.2] point out that surface  
296 wave/internal wave coupling has been previously regarded primarily as a *sink* of internal  
297 wave energy. One might conjecture that swell could interact efficiently with a shallow  
298 stratification, along the lines of the well-known ‘dead water’ phenomenon of ship drag from  
299 interfacial waves in fjords [see e.g. *Gill*, 1982, Section 6.2]. However, *Polzin and L'vov*  
300 [2011] state, “rapid transfers of energy from a narrowband ocean swell to the internal  
301 wavefield are possible ... and remain to be quantified,” thus it appears that little is known  
302 about this possible wave source.

#### 4. Conclusions

303 Microstructure measurements from the Labrador Sea in May 2010 show a 10 m am-  
304 plitude internal wave train, with a 50 minute period, associated with elevated levels of  
305 turbulence in the surface mixed layer. The breaking internal wave event observed here  
306 increased the average turbulent dissipation rate in the pycnocline from  $7 \times 10^{-10} \text{ m}^2\text{s}^{-3}$  to  
307  $1 \times 10^{-8} \text{ m}^2\text{s}^{-3}$  and the average vertical eddy diffusivity from  $6 \times 10^{-6} \text{ m}^2\text{s}^{-1}$  to  $1 \times 10^{-4}$   
308  $\text{m}^2\text{s}^{-1}$ , with maximum values of  $2 \times 10^{-7} \text{ m}^2\text{s}^{-3}$  and  $6 \times 10^{-4} \text{ m}^2\text{s}^{-1}$  respectively. The eddy

309 diffusivity is higher than the values estimated through internal wave parameterisations  
310 by *Lauderdale et al.* [2008] near the bottom of the central basin of the Labrador Sea and  
311 on par with those estimated in the same way by *Walter et al.* [2005] at mid-depths in the  
312 central basin, but in the latter study the elevated mixing was due to reduced stratification,  
313 not enhanced turbulence.

314 While observational limitations prevent us from identifying the exact breaking mecha-  
315 nism observed here, interaction of the internal wave and background shear appears to be  
316 the most likely candidate. These results appear to be the direct evidence that breaking  
317 internal waves may play a role in driving near-surface turbulence. This raises the ques-  
318 tion as to how widespread such features may be. With a profile interval of 5–10 minutes,  
319 ASIP’s sampling rate is far higher than most other observational platforms, and its ability  
320 to profile upward to the sea surface is rare among microstructure instruments. Thus, even  
321 if such the breaking of internal waves at the mixed layer base are a commonplace occur-  
322 rence, few of the standard oceanographic measurement platforms would be able to capture  
323 them. It is conceivable that such phenomenon are widespread but been overlooked due  
324 to observational limitations. Another possibility is that the occurrence of such a wave in  
325 the Labrador Sea is a particular consequence of the highly surface-intensified nature of  
326 the stratification in this region. If so, one may expect to find such waves in other regions  
327 with similar surface-intensified stratifications, such as the Gulf of Lion, Greenland Sea,  
328 and other regions of deep convection [*Marshall and Schott*, 1999]. Additional work is  
329 necessary to determine whether shallow dissipation by such waves may play a significant  
330 role in the restratification process following deep convection.

331 Processes such as mixing due to near-inertial shear [*Dohan and Sutherland*, 2003], and  
332 high-frequency wave generation associated with Langmuir cells [*Polton et al.*, 2008], are  
333 already recognized as important processes in near-surface mixing. Recently, *Belcher et al.*  
334 [2012] showed that current surface turbulence parameterizations lead to biases in coupled  
335 climate models, a fact that they argued was due in part to the failure to correctly account  
336 for turbulence associated with Langmuir circulations. They concluded that there is a  
337 pressing need for more microstructure measurements within the ocean surface boundary  
338 layer. Our results support this assessment, by indicating that there may still be other  
339 processes contributing significantly to near-surface mixing that have yet to be accounted  
340 for.

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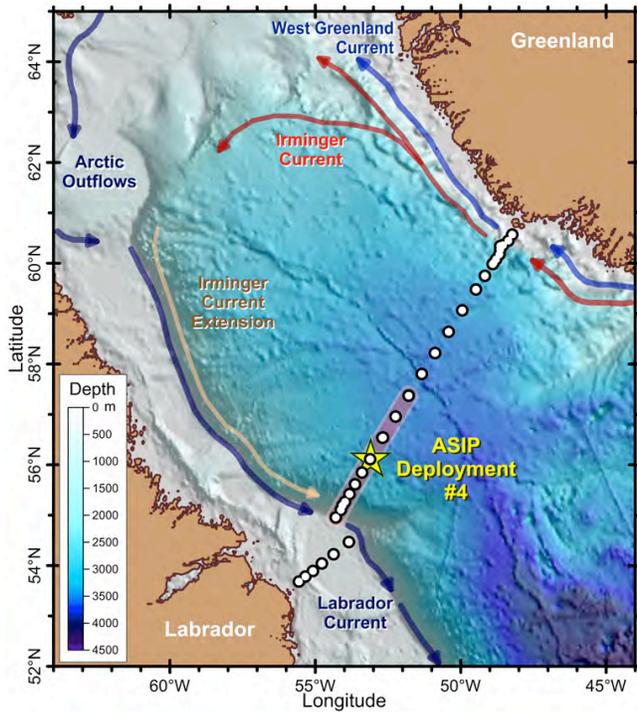
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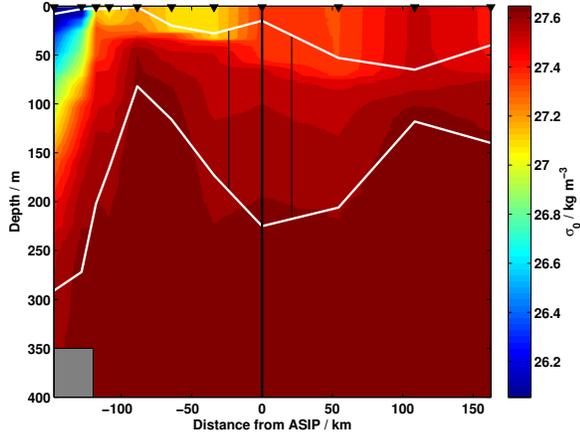
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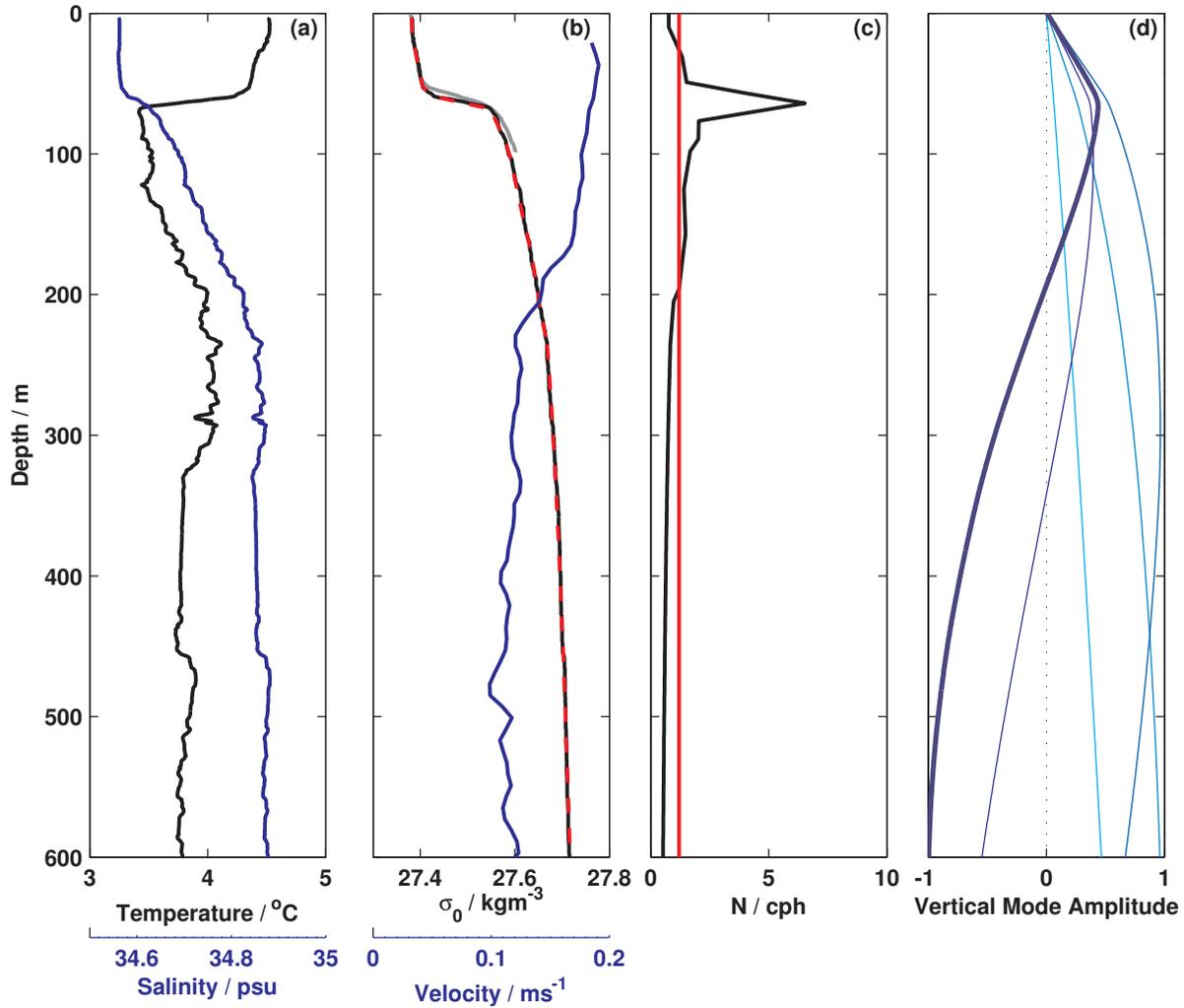
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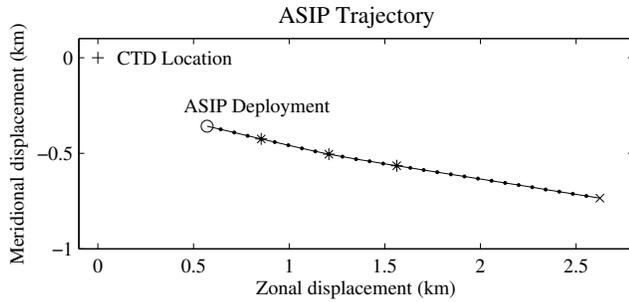
**Figure 1.** Bathymetric map of the Labrador Sea with primary currents labeled. The position of the CTD casts from the 2010 survey of the AR7W line are indicated by filled white circles. The profiles used to create the transect in Figure 2 are highlighted in purple.



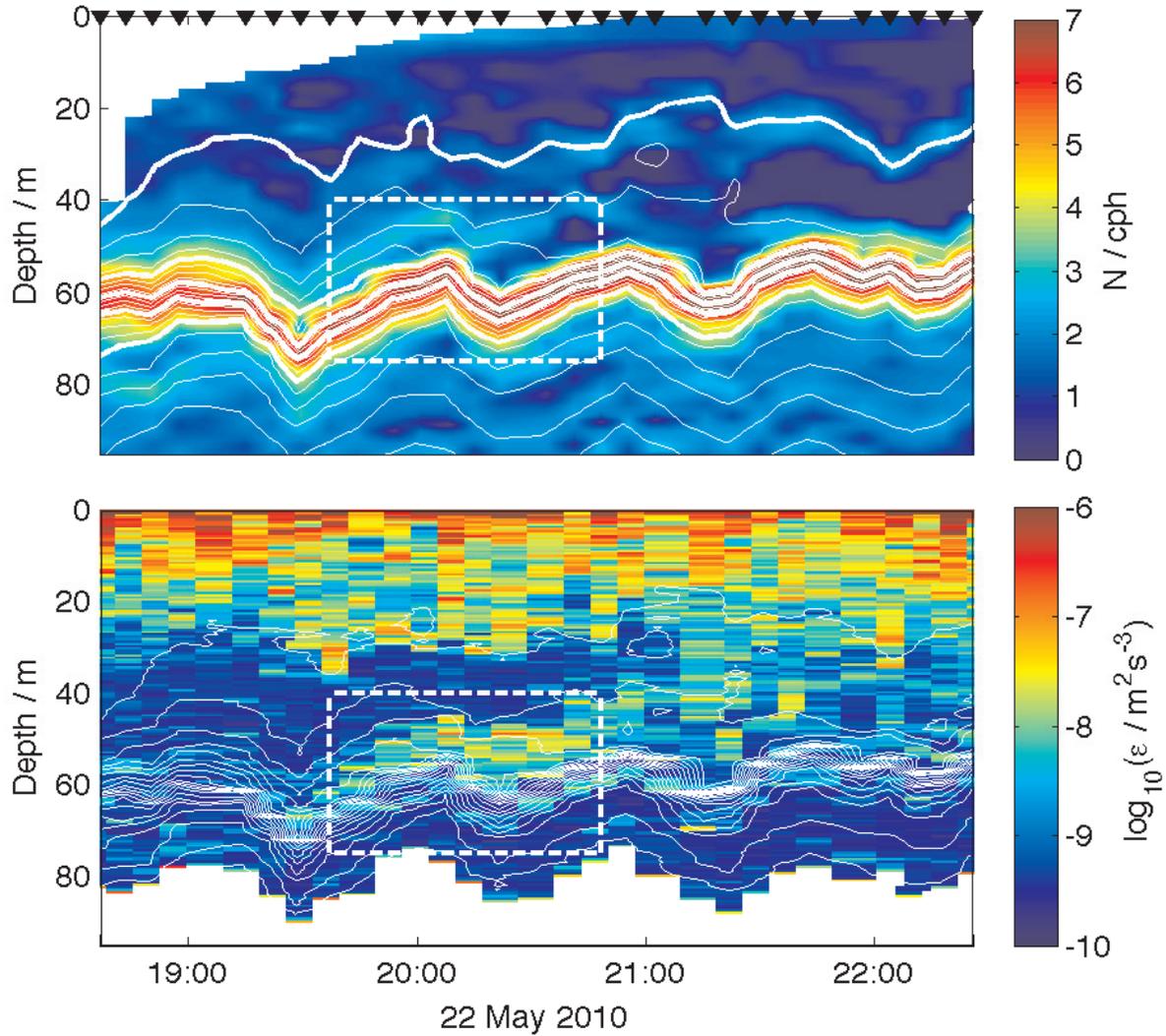
**Figure 2.** Density section across the upper 350 m of the western portion of the AR7W transect, corresponding the portion highlighted in purple in Figure 1. The black triangles indicate the location of the CTD casts. The envelope of  $N$  equal to the frequency of the observed waves is indicated by white lines, which therefore delimit the upper and lower boundaries of the ray paths. The thick vertical black line indicates the approximate position of the ASIP measurements, and the distances on the  $x$ -axis are computed relative to this point. The thin black line is the horizontal distance that a wave of this frequency would travel in 100 reflections off the boundaries of the  $N$  envelope. The grey box in the lower left is due to a decrease in bottom depth approaching the continental slope.



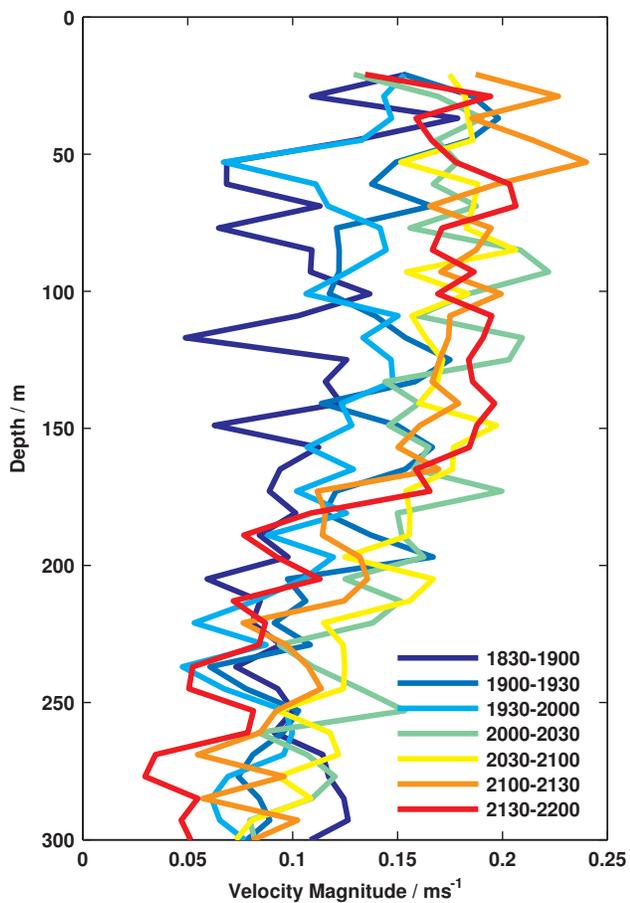
**Figure 3.** (a) Temperature and salinity profiles from the shipboard CTD cast concurrent with the ASIP measurements. (b) Density profile from the CTD, together with velocity magnitude from the shipboard ADCP. The piecewise linear fit to the density profile is shown as a red dashed line overlying the original data. The mean ASIP density profile is shown in grey. (c) Buoyancy frequency profile. The vertical red line indicates the estimated frequency— $0.002 \text{ rad s}^{-1}$ , or a 50 minute period—of the internal wave. (d) The upper 600 m of the vertical structure functions for the first five vertical modes, with mode one in the lightest color and proceeding to mode five in the darkest color. Mode five is indicated as a heavy line.



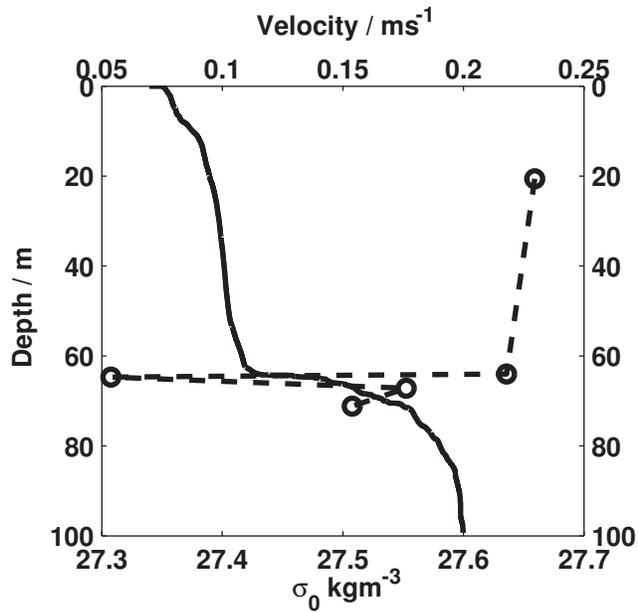
**Figure 4.** The estimated trajectory of ASIP relative to the CTD cast, indicated by a “+”, during the deployment on 22 May 2010. ASIP was deployed from a small boat about a kilometer away from the *Hudson* at the location marked by a circle. Note that only the three points marked by asterisks are based on valid GPS fixes; the dots are estimated locations based on interpolating or extrapolating the velocities from the valid points.



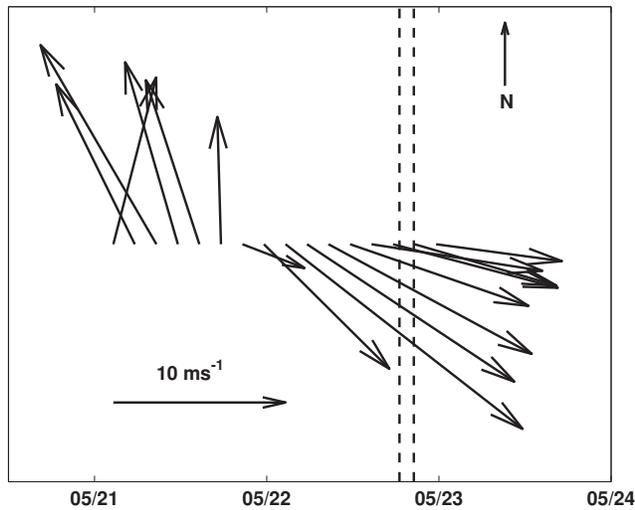
**Figure 5.** Four-hour records of: (a) Buoyancy frequency  $N$  overlaid by the isopycnals of the smoothed density profiles. (b) Dissipation rate of turbulent kinetic energy  $\epsilon$  overlaid by the isopycnals from the unsmoothed density profiles. The time of the profiles is marked by the black triangles. All isopycnals (white lines) are spaced at  $0.01 \text{ kg m}^{-3}$  intervals. The isopycnals used to determine the profile in Figure 7 are highlighted. The elevated dissipation due to the breaking internal wave is demarcated by the dashed box.



**Figure 6.** 30 minute averages of the velocity magnitude measured by the shipboard ADCP.



**Figure 7.** Sorted density profile (solid line) from Profile 10 (measured at 19:43) and inferred velocity profile (dashed line) following the method set forth in *Moum et al.* [2003].



**Figure 8.** Wind speed and direction from ECMWF at the position of ASIP before and during the deployment.